

the Structure and Evolution of Single Stars

AN INTRODUCTION

James MacDonald

Structure Evolution Single Stars Introduction

**International Astronomical Union.
Symposium**



Structure Evolution Single Stars Introduction:

Theory of Rotating Stars. (PSA-1), Volume 1 Jean-Louis Tassoul, 2015-03-08 Ever since the first observations of sunspots in the early seventeenth century stellar rotation has been a major topic in astronomy and astrophysics Jean Louis Tassoul synthesizes a large number of theoretical investigations on rotating stars Drawing upon his own research Professor Tassoul also carefully critiques various competing ideas In the first three chapters the author provides a short historical sketch of stellar rotation the main observational data on the Sun and other stars on which the subsequent theory is based and the basic Newtonian hydrodynamics used to study rotating stars Following a discussion of some general mechanical properties of stars in a state of permanent rotation he reviews the main techniques for determining the structure of a rotating star and its stability with respect to infinitesimal disturbances Since the actual distribution of angular momentum within stars is still unknown Professor Tassoul considers various models of angular momentum as well as of meridional circulation He devotes the rest of his study to the problems concerning various groups of stars and stages in stellar evolution Originally published in 1979 The Princeton Legacy Library uses the latest print on demand technology to again make available previously out of print books from the distinguished backlist of Princeton University Press These editions preserve the original texts of these important books while presenting them in durable paperback and hardcover editions The goal of the Princeton Legacy Library is to vastly increase access to the rich scholarly heritage found in the thousands of books published by Princeton University Press since its founding in 1905

An Introduction to Close Binary Stars R. W. Hilditch, 2001-03-12 Binary systems of stars are as common as single stars They are of fundamental importance because they allow stellar masses radii and luminosities to be measured directly and explain a host of diverse and energetic phenomena including X ray binaries cataclysmic variables novae symbiotic stars and some types of supernovae This 2001 book was the first to provide a pedagogical and comprehensive introduction to binary stars It combines theory and observations at all wavelengths to develop a unified understanding of binaries of all categories It comprehensively reviews methods for calculating orbits the Roche model ideas about mass exchange and loss methods for analysing light curves the masses and dimensions of different binary systems and imaging the surfaces of stars and accretion structures This book provides a thorough introduction to the subject for advanced undergraduate and graduate students Researchers will also find this to be an authoritative reference

Physics of Binary Star Evolution Thomas M. Tauris, Edward P.J. van den Heuvel, 2023-03-28 A graduate level textbook on the astrophysics of binary star systems and their evolution Physics of Binary Star Evolution is an up to date textbook on the astrophysics and evolution of binary star systems Theoretical astrophysicists Thomas Tauris and Edward van den Heuvel cover a wide range of phenomena and processes including mass transfer and ejection common envelopes novae and supernovae X ray binaries millisecond radio pulsars and gravitational wave GW sources and their links to stellar evolution The authors walk through the observed properties and evolution of different types of binaries with special emphasis on those

containing compact objects neutron stars black holes and white dwarfs Attention is given to the formation mechanisms of GW sources merging double neutron stars and black holes as well as ultra compact GW binaries hosting white dwarfs and to the progenitors of these sources and how they are observed with radio telescopes X ray satellites and GW detectors LIGO Virgo KAGRA Einstein Telescope Cosmic Explorer and LISA Supported by illustrations equations and exercises Physics of Binary Star Evolution combines theory and observations to guide readers through the wonders of a field that will play a central role in modern astrophysics for decades to come 465 equations 47 tables and 350 figures More than 80 exercises analytical numerical and computational Over 2 500 extensive up to date references

Structure and Evolution of Single Stars James MacDonald, 2015-12-01 Structure and Evolution of Single Stars An introduction is intended for upper level undergraduates and beginning graduates with a background in physics Following a brief overview of the background observational material the basic equations describing the structure and evolution of single stars are derived The relevant physical processes which include the equation of state opacity nuclear reactions and neutrino losses are then reviewed Subsequent chapters describe the evolution of low mass stars from formation to the final white dwarf phase The final chapter deals with the evolution of massive stars

Nuclear Physics of Stars Christian Iliadis, 2015-03-19 Most elements are synthesized or cooked by thermonuclear reactions in stars The newly formed elements are released into the interstellar medium during a star's lifetime and are subsequently incorporated into a new generation of stars into the planets that form around the stars and into the life forms that originate on the planets Moreover the energy we depend on for life originates from nuclear reactions that occur at the center of the Sun Synthesis of the elements and nuclear energy production in stars are the topics of nuclear astrophysics which is the subject of this book It presents nuclear structure and reactions thermonuclear reaction rates experimental nuclear methods and nucleosynthesis in detail These topics are discussed in a coherent way enabling the reader to grasp their interconnections intuitively The book serves both as a textbook for advanced undergraduate and graduate students with worked examples and end of chapter exercises but also as a reference book for use by researchers working in the field of nuclear astrophysics

Triggered Star Formation in a Turbulent Interstellar Medium (IAU S237) International Astronomical Union. Symposium, 2007-05-28 New stars form in the dense turbulent gas clouds of galaxies and the formation of these clouds is the subject of the IAU S237 This book is the most up to date review of all aspects of cloud and star formation and one of the few compendiums available on ISM turbulence

Convection in Astrophysics (IAU S239) International Astronomical Union. Symposium, 2007-06-07 Convection is ubiquitous throughout the Universe and during the last three decades it has become the largest factor of uncertainty in theoretical models of stars and in the interpretation of observations on the basis of such models Recently numerical simulations of convection have dramatically improved in their potential to take into account both the large scale properties of the flow itself and the microphysical properties of the fluid Observations have become accurate enough to provide stringent tests for both numerical simulations and models of

convection IAU S239 was held to further understanding of convection bringing together leading researchers in solar and stellar physics the physics of planets and of accretion disks With reviews research contributions and detailed recordings of plenary discussions this book is a valuable resource for professional astronomers and graduate students interested in the interdisciplinary study of one of the key physical processes in astrophysics The Brightest Binaries D. Vanbeveren,W. van Rensbergen,C. de Loore,2001-11-30 The evolution of galaxies is governed mainly by the evolution of massive stars whereas the evolution of a massive star depends primarily on its mass chemical composition and on whether or not the star is a Single object or a binary component To study the evolution of galaxies it is therefore essential to know how stellar masses are distributed at birth how many stars are formed in binaries and what the mass ratio and orbital period distribution of binaries look like Massive stars are intrinsically the brightest stars so that it may be possible to discover their properties in distant groups prOvided that large telescopes can be used for basic stellar observations However until now the observations of massive stars have been reasonably complete only for a small region of our own Galaxy 3 kpc from the Sun One hopes that the conclusions resulting from these observations hold for the whole Galaxy for the whole cosmos With The Brightest Stars of De Jager 1980 in mind the present monograph is an addendum and an update in which we discuss the observations of The Brightest Binaries in the framework of stellar evolution A small or intermediate mass star close to the Sun may look brighter than a massive one far away However within volume limited star samples the massive stars are on average also the brightest ones In the present monograph similarly as in the work of De Jager bright means massive The book consists of four main chapters *Cataclysmic Variables and Related Objects: pp.i-xc, 1-342* Margherita Hack,Constanze La Dous,1993 **The Art of Modeling Stars in the 21st Century (IAU S252)** International Astronomical Union. Symposium,2008-12-18 Stellar models are the very basic building blocks with which we build up our knowledge of the Universe New numerical experiments are heralding a new level of sophistication in our ability to model and understand how stars work This volume provides an overview and the most recent advances in modeling of stellar structure and evolution Modeling of stars relies on our understanding of the detailed physical processes happening in stars and the most recent observations of stars made by modern large telescopes and current high technologies IAU Symposium 252 presents the most recent developments in five key areas including improvements of the physical ingredients of stellar models the evolution of low and intermediate mass stars the evolution of massive stars close binary evolution and stellar physics in the era of very large telescopes This overview of stellar research is at a level suitable for research astronomers and graduate students

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